## Photon equation dimensional:

For the photon equation we assume the hypothesis that the instant value of light quanta H(r.z,t) is proportional to light intensity and satisfies the equation:

$$\frac{R}{r}\Delta H(r,\theta,z,t) = \frac{1}{c^2} \frac{\partial^2 H(r,\theta,z,t)}{\partial t^2}$$
(1)

Where R is a fixed distance, the same as in the graviton equation. With the initial condition for photon H(z=R) = 0 with z cylindrical symmetry so the initial value is null for a cylindrical surface of a cylinder of radius R, a constant of distance equal to  $R = 3.567 \cdot 10^{22}$  m. In cylindrical coordinates with z-axis symmetry the Laplacian for the photon is considered that be invariant of  $\theta$  (z-axis symmetry):

$$\Delta H_2(r,z) = \frac{1}{r} \frac{\partial}{\partial r} \left( r \frac{\partial (H_2(r,z))}{\partial r} \right) + \frac{\partial^2 H_2(r,z)}{\partial z^2}$$

$$\begin{pmatrix} \partial (H_2(r,z)) u(t) \\ \partial z^2 H_2(r,z) u(t) \end{pmatrix} = \frac{\partial^2 H_2(r,z) u(t)}{\partial z^2 H_2(r,z) u(t)}$$
(2)

$$\frac{R}{r}\left(\frac{1}{r}\frac{\partial}{\partial r}\left(r\frac{\partial(H_2(r,z))u(t)}{\partial r}\right) + \frac{\partial^2 H_2(r,z)u(t)}{\partial z^2}\right) = \frac{1}{c^2}\frac{\partial^2 H_2(r,z)u(t)}{\partial t^2}$$
(3)  
With the method of separation of variables  $H_2(r,t)$  not null  $H(r,z,t) = H_1(r,z)$ ,  $u(t)$  (4)

With the method of separation of variables  $H_2(r,t)$  not null  $H(r,z,t) = H_2(r,z) \cdot u(t)$  (4)

$$\frac{R}{rH_2(r,z)} \left( \frac{1}{r} \frac{\partial}{\partial r} \left( r \frac{\partial (H_2(r,z))u(t)}{\partial r} \right) + \frac{\partial^2 H_2(r,z)u(t)}{\partial z^2} \right) = \frac{1}{c^2} \frac{\partial^2 u(t)}{\partial t^2}$$

(8)

Thus we split it into two equations:  $\frac{\kappa}{r} \left( \frac{1}{r} \frac{\partial}{\partial r} \left( r \frac{\partial (H_2(r,2))}{\partial r} \right) + \frac{\partial (H_2(r,2))}{\partial z^2} \right) = b^2 H_2(r,z)$  (5)

And 
$$\frac{1}{c^2} \frac{\partial^2 u(t)}{\partial t^2} = b^2$$
, (6)

$$\left(\frac{1}{r}\frac{\partial}{\partial r}\left(r\frac{\partial(H_2(r,z))}{\partial r}\right) + \frac{\partial^2 H_2(r,z)}{\partial z^2}\right) - \frac{rb^2 H_2(r,z)}{R} = 0$$
(7)

With the initial value at the limit H<sub>3</sub>(1/R)=R and H<sub>3</sub>(R)=0;  $A^{2}H_{1}(r)H_{1}(r)H_{2}(r)H_{2}(r)H_{2}(r)H_{3}(r)$ 

$$\left(\frac{1}{r}\frac{\partial}{\partial r}\left(r\frac{\partial(H_3(r)H_4(z))}{\partial r}\right) + \frac{\partial^2 H_3(r)H_4(z)}{\partial z^2}\right) - \frac{rb^2 H_3(r)H_4(z)}{R} = 0$$

$$\frac{1}{r}\frac{\partial}{\partial r}\left(r\frac{\partial(H_3(r)H_4(z))}{\partial r}\right) - \frac{rb^2H_3(r)H_4(z)}{R} = -\frac{\partial^2H_3(r)H_4(z)}{\partial z^2}$$
(9)

$$\frac{1}{rH_3(r)}\frac{\partial}{\partial r}\left(r\frac{\partial(H_3(r))}{\partial r}\right) - \frac{rb^2}{R} = -\frac{\partial^2 H_4(z)}{H_4(z)\partial z^2} \tag{10}$$

$$\frac{1}{rh^2}\frac{\partial}{\partial r}\left(\frac{\partial(H_3(r))}{\partial r}\right) - \frac{rb^2}{R^2} = -\frac{\partial^2 H_4(z)}{H_4(z)\partial z^2} \tag{11}$$

$$\frac{1}{rH_3(r)}\frac{\partial}{\partial r}\left(r\frac{\partial(H_3(r))}{\partial r}\right) - \frac{rb^2}{R} = -\nu^2 \tag{11}$$

Hence 
$$\frac{1}{r}\frac{\partial}{\partial r}\left(r\frac{\partial(H_3(r))}{\partial r}\right) + H_3(r)\left(-\frac{rb^2}{R} + \nu^2\right) = 0$$
(12)
$$\frac{\partial^2(H_3(r))}{\partial r} + \frac{\partial(H_3(r))}{\partial r} + \frac{c^2b^2}{R} + \frac{2}{r^2} + \frac{2}{$$

$$r\frac{\sigma(H_3(r))}{\partial r^2} + \frac{\sigma(H_3(r))}{\partial r} - (\frac{r}{R} - v^2 r)H_3(r) = 0$$
(13)

And 
$$\frac{\partial^2 H_4(z)}{\partial z^2} = -\nu^2 H_4(z)$$
 (14)

With sinusoidal solution  $H_4(z) = \sin(\nu z + \varphi) = \sin\left(\frac{E}{h}z + \varphi\right)$  where  $\nu$  is positive constant(15) Considering the photon energy E=h· $\nu$  and we have the wave pulsation of  $\nu \cdot z$  thus  $\frac{E}{h} = \nu$  (16) So we have:  $H_4(z) = \sin\left(\frac{E}{h}z + \varphi\right)$  (17)

with dependence on z-axis:  $H_4(z) = \sin(\frac{E}{h}z + \varphi)$  where h = Planck constant, E photon energy, and  $\varphi$  the phase with H<sub>4</sub>(z) and u(t) a dimensional function;

The other initial value for equation (7) is referring to the initial velocity of the photon which is c the speed of light all the time hence  $d = 1 \cdot [L^{-3/2}]$  for the dimensional reason we have:

$$\frac{df(r,z,t)}{dt} = \frac{\partial[u(t)H_3(r)H_4(z)]}{\partial t} + \frac{\partial[u(t)H_3(r)H_4(z)]}{\partial r} \cdot \frac{\partial r}{\partial t} + \frac{\partial[u(t)H_3(r)H_4(z)]}{\partial z} \cdot \frac{\partial z}{\partial t} = c \cdot d$$
(19)  
$$c \cdot d = \frac{\partial[u(t)H_3(r)H_4(z)]}{\partial ct} + \frac{\partial[u(t)H_3(r)H_4(z)]}{\partial r} + \frac{\partial[u(t)H_3(r)H_4(z)]}{\partial z} \cdot \frac{\partial z}{\partial t}$$
(20)

Divide by "c" both members of the equation:

$$\frac{\partial [u(t)H_3(r)H_4(z)]}{\partial ct} + \frac{\partial [u(t)H_3(r)H_4(z)]}{c\partial r} \cdot \frac{\partial r}{\partial t} + \frac{\partial [u(t)H_3(r)H_4(z)]}{c\partial z} \cdot \frac{\partial z}{\partial r} \frac{\partial r}{\partial t} = d$$
(21)

But 
$$\frac{\partial z}{\partial r} = 0$$
 and  $\frac{\partial r}{\partial t} = c$ , so:  $\frac{\partial [u(t)H_3(r)H_4(z)]}{\partial ct} + \frac{\partial [u(t)H_3(r)H_4(z)]}{\partial r} = d$  (22)

$$H_4(z)H_3(r)\frac{\partial u(t)}{\partial ct} + u(t)H_4(z)\frac{\partial [H_3(r)]}{\partial r} = d \qquad \text{for } t = 0 \text{ and } d = 1 \cdot [L^{-3/2}]$$
(23)

$$H_3(r)\frac{\partial u(t)}{\partial ct} + u(t)\frac{\partial [H_3(r)]}{\partial r} = \frac{1}{H_4(z)}$$
(24)

Thus as above according to (10-13) we have to solve the equation:

$$r\frac{d^{2}(H_{3}(r))}{dr^{2}} + \frac{d(H_{3}(r))}{dr} + \left(-\frac{r^{2}b^{2}}{R} + \left(\frac{E}{h}\right)^{2}r\right)H_{3}(r) = 0$$
(25)  
With the initial value at the limit H<sub>3</sub>(1/R)=R and H<sub>3</sub>(R)=0;

$$\frac{d^2(H_3(r))}{dr^2} + \frac{d(H_3(r))}{r \cdot dr} + \left(-\frac{r \cdot b^2}{R} + \left(\frac{E}{h}\right)^2\right) H_3(r) = 0$$
(26)

The equation r y'' + y' + r y = 0 cannot be solved in terms of elementary functions. Where:  $v = \frac{E}{h}$ 

But for R=3.567 · 10<sup>22</sup> m and wavelength of 300nm frequency of 10<sup>15</sup> Hz and r between 1/R < r < R we can neglect them  $\frac{rb^2}{R}$  and we have approximately the equation:  $\frac{d^2(r \cdot H(r))}{dr^2} + \nu^2 H(r) = 0$  equation (27);

With boundary condition can be solved with Maple bc:= H<sub>3</sub>(R)=0,  $\frac{d(H(\frac{1}{R}))}{dr} = R$  Thus the solution is H<sub>3</sub>(r)

 $H(r, z, t) = -H_3(r) \left( \sin\left(\frac{E}{h}z + \varphi\right) \right) u(t)$ (29) From (23) we have  $H_4(z)H_3(r)\frac{\partial u(t)}{\partial ct} + u(t)H_4(z)\frac{\partial [H_3(r)]}{\partial r} = 1$  (d unitary) where we know all the functions except u(t). Thus equation (24):  $H_3(r)\frac{\partial u(t)}{\partial ct} + u(t)\frac{\partial [H_3(r)]}{\partial r} = \frac{1}{H_4(z)}$ 

Thus, we solve the equation, and we have EOWF hence Photon => we divide H<sub>3</sub>(r)/r with d unitary of dimension [L<sup>-3/2</sup>] the final solution approximative of the photon onduscular wave function H<sub>3</sub>(r)H<sub>4</sub>(z)·u(t)/r of dim [L<sup>-3/2</sup>] with C2 normed constant and \_C1 as in file <u>http://michaelvio.orgfree.com/PhotonA.pdf</u> Viktor T. Toth: *"However, an important milestone was the development of the Proca theory (named after Romanian physicist Alexandru Proca) in the late 1930s. This was the theory of massive vector mesons in nuclear physics; the corresponding classical theory, known nowadays as Maxwell—Proca theory, is the massive alternative of Maxwell's theory. " Does the photon have rest mass <>0? it could be a rough estimation of the rest mass is ~1.41 \cdot 10<sup>-9</sup> eV/c<sup>2</sup>. Let's take o photon of wavelength equal to 620nm thus an energy of 0.5eV. A magnetron thus a pion \pi^+, \pi^- has a rest mass of around 0.00678eV/c<sup>2</sup> see the <u>http://michaelvio.orgfree.com/Magnet.pdf</u>* 

The energy of a magnetron of f=100Mhz is E=hm·f so also 0.5eV, but hm/h equals  $4.79461495 \cdot 10^6$  thus rest mass of 620nm photon should be around ~1 - 2  $\cdot 10^{-9}$  eV/c<sup>2</sup>. <u>http://michaelvio.orgfree.com/grating.pdf</u> Thus, we have for massless photon equation (30):

$$Photon := \frac{1}{r} \left( \frac{bR \operatorname{BesselY}(0, kR) \operatorname{BesselJ}(0, kr)}{k \left( \operatorname{BesselY}(1, \frac{ka}{R}) \operatorname{BesselJ}(0, kR) - \operatorname{BesselJ}(1, \frac{ka}{R}) \operatorname{BesselY}(0, kR) \right)} \right)$$

$$-\frac{\operatorname{BesselJ}(0, kR) bR\operatorname{BesselY}(0, kr)}{k\left(\operatorname{BesselY}\left(1, \frac{ka}{R}\right)\operatorname{BesselJ}(0, kR) - \operatorname{BesselJ}\left(1, \frac{ka}{R}\right)\operatorname{BesselY}(0, kR)\right)}\right)$$
$$\left(\frac{k}{C2\sin(\nu z + \phi)\left(\frac{d^2}{dr^2}H(r)\right)} + e^{-\frac{\left(\frac{d^2}{dr^2}H(r)\right)c^2t}{H(r)}}CI\right)$$

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